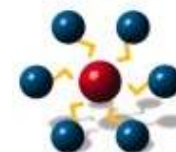
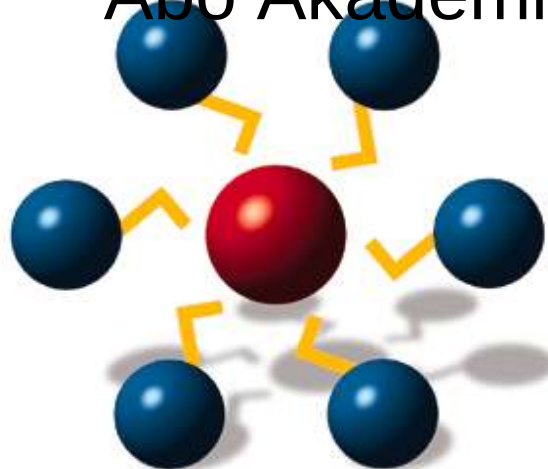


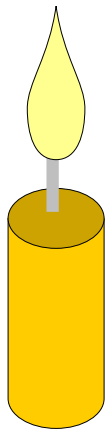
Spektrometer

Matti Hotokka
Institutionen för Fysikalisk Kemi
Åbo Akademi

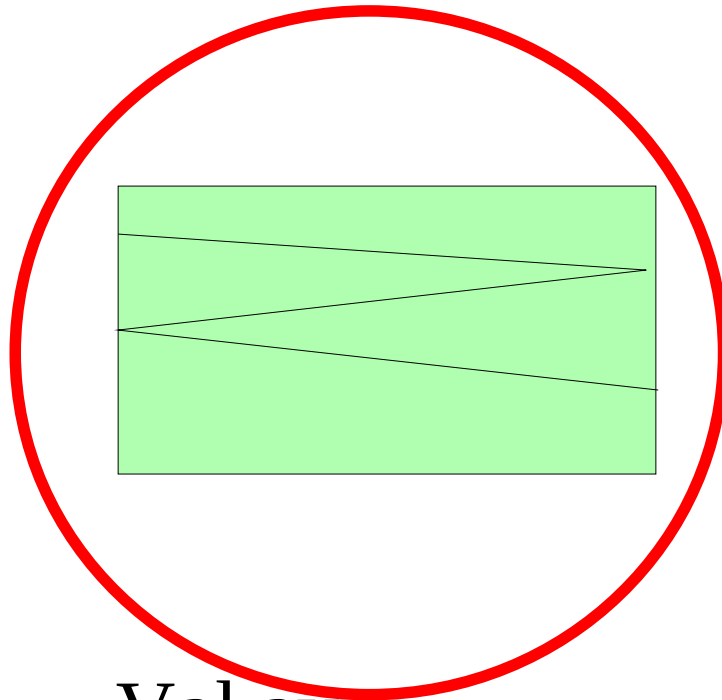


Spektrometer schematiskt

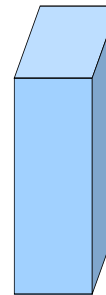
Interferometer



Källa



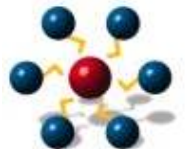
Val av
våglängd



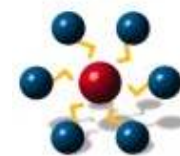
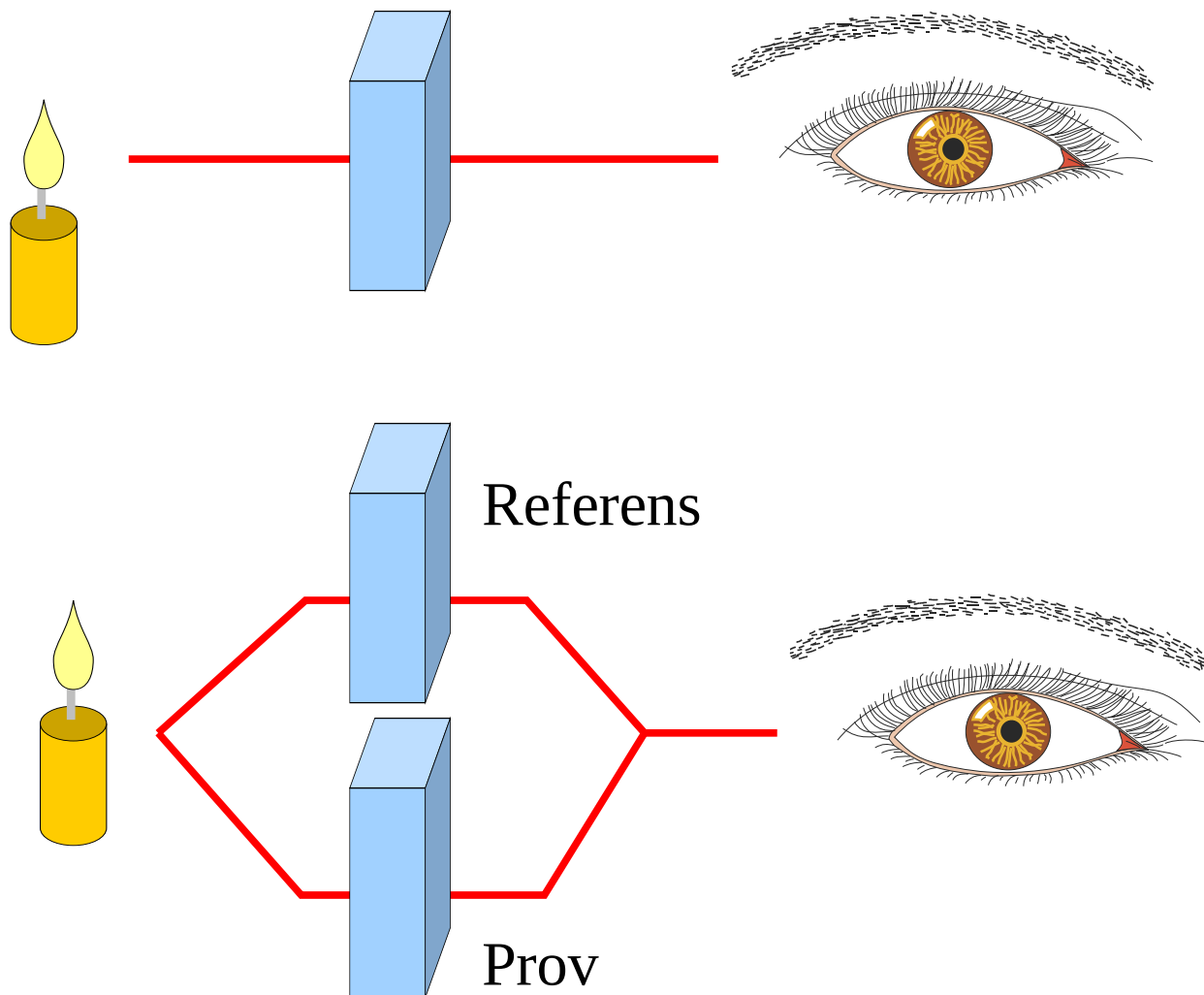
Prov



Detektor

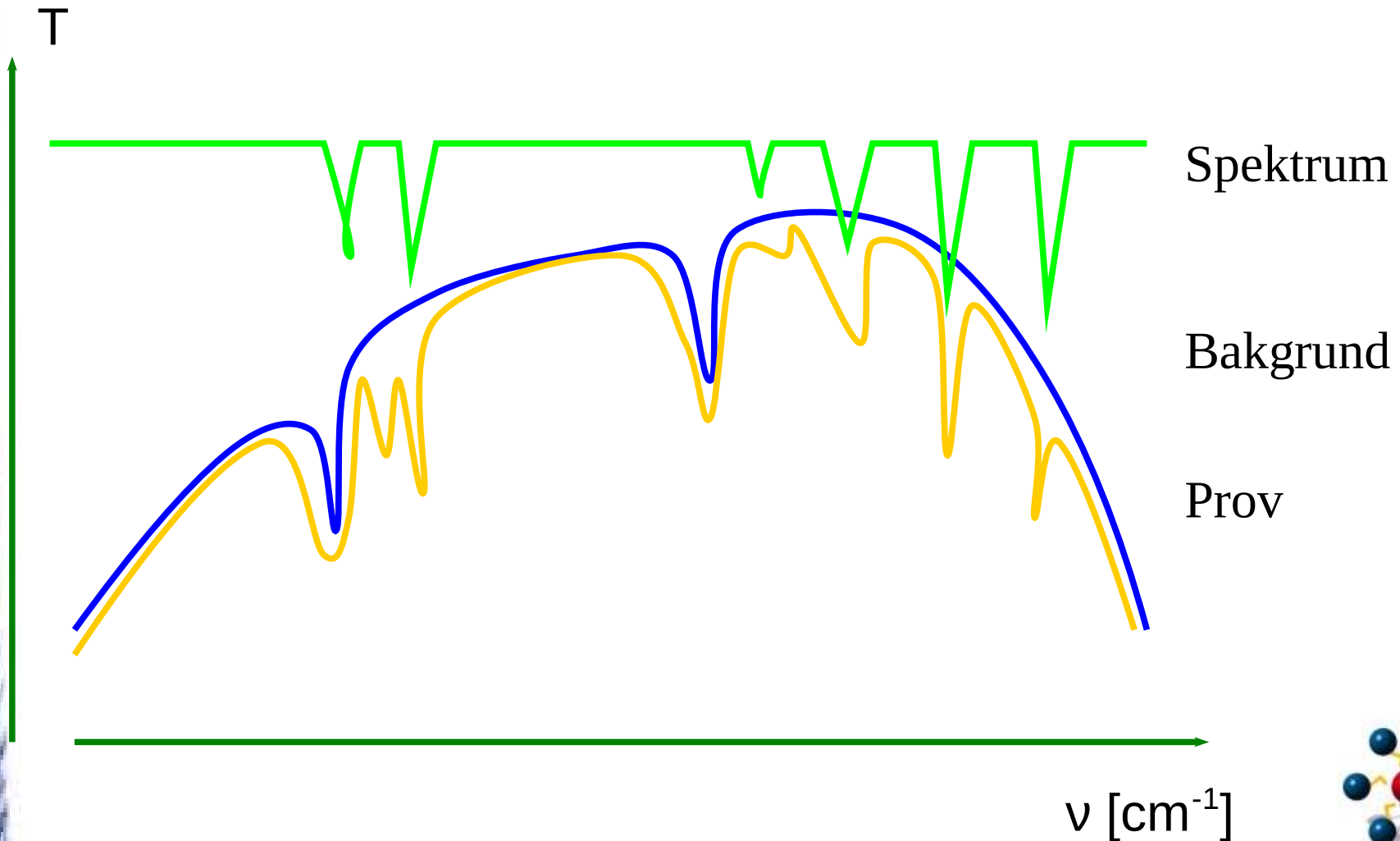


Enstråleapparat



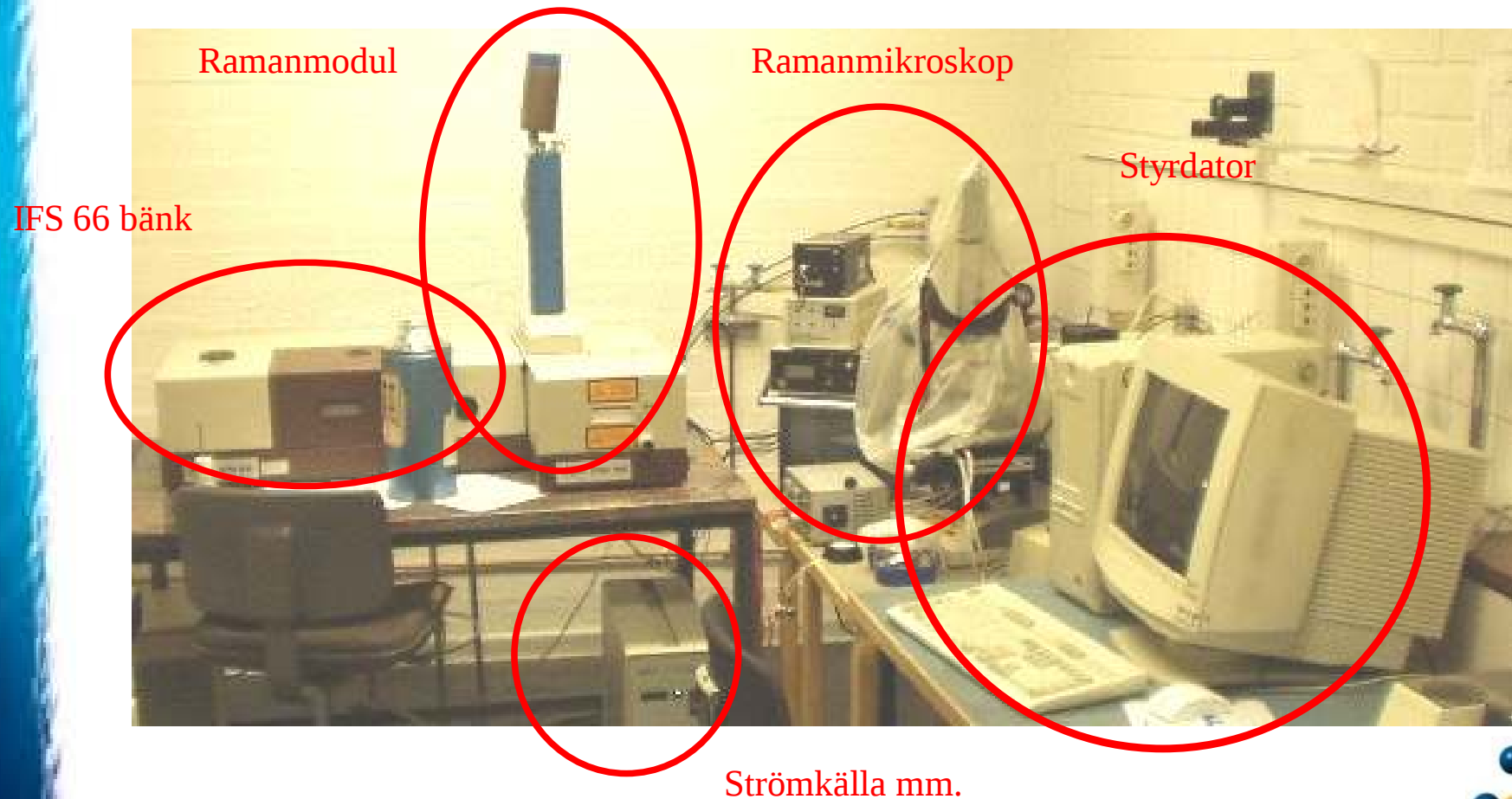
Enstråleapparat

Spektret fås genom att subtrahera



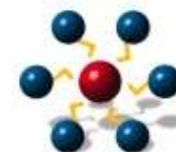
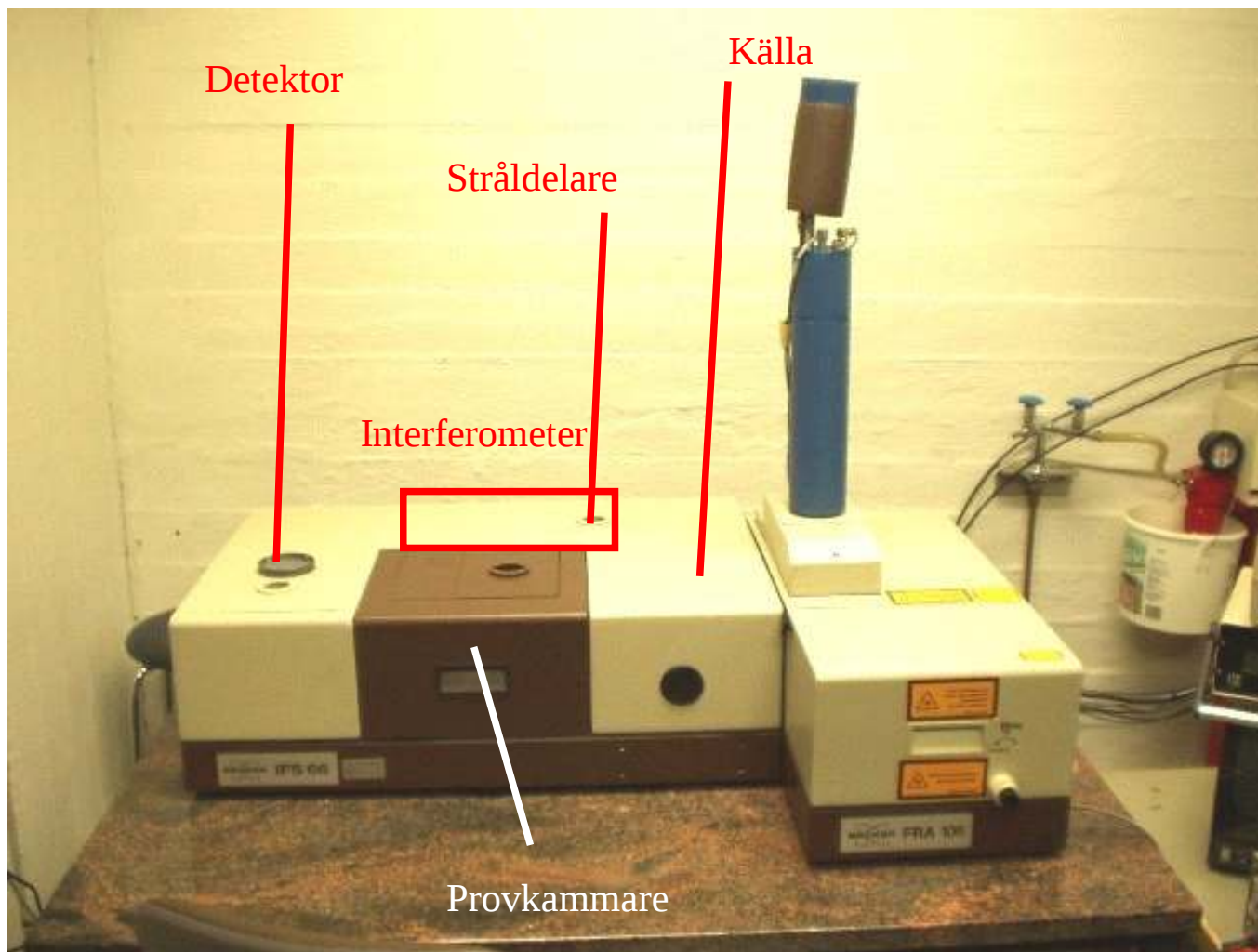
Exempel på apparat

Bruker IFS66



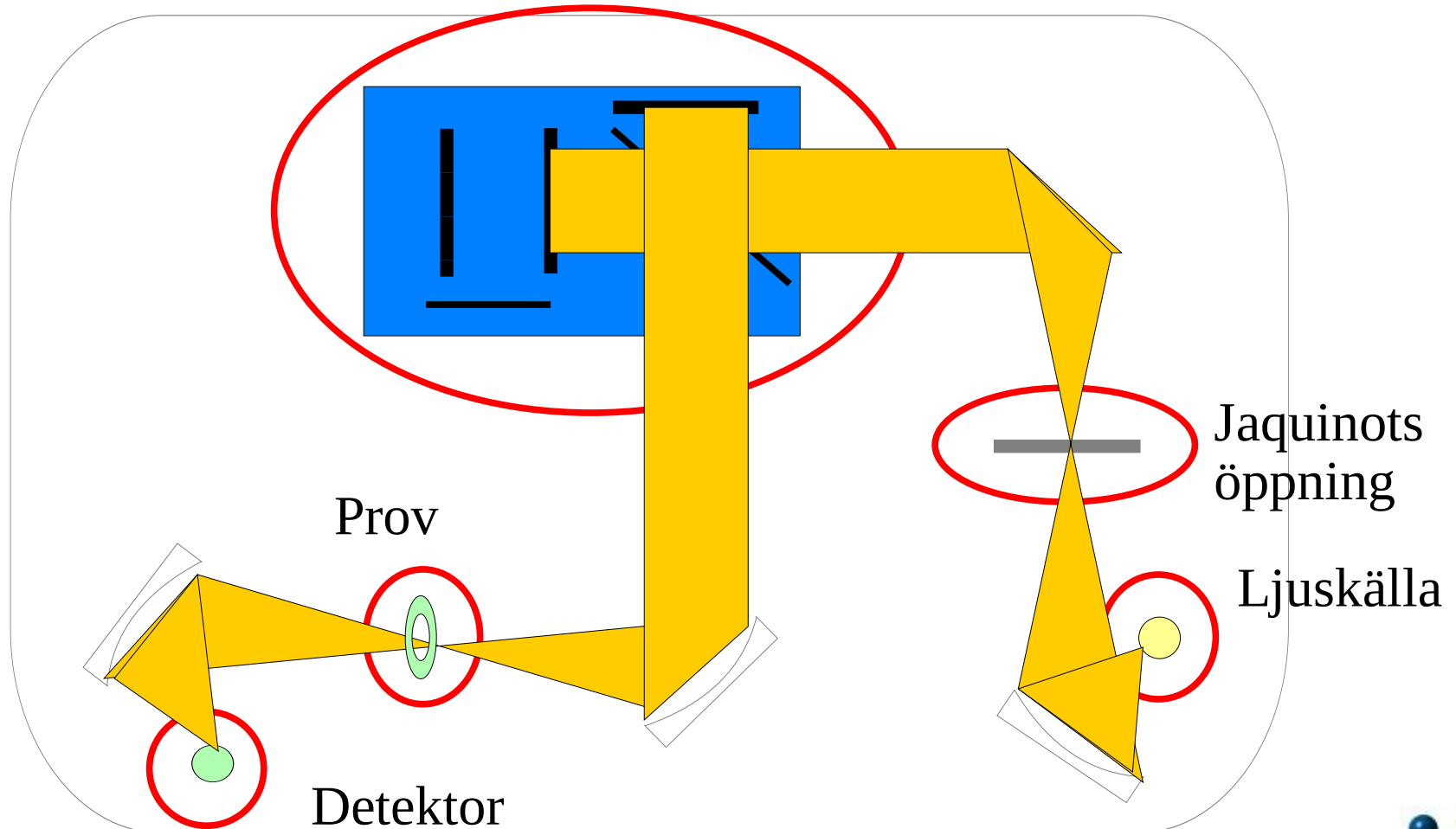
Exempelapparat

Bruker IFS66



Bruker IFS66

Interferometer



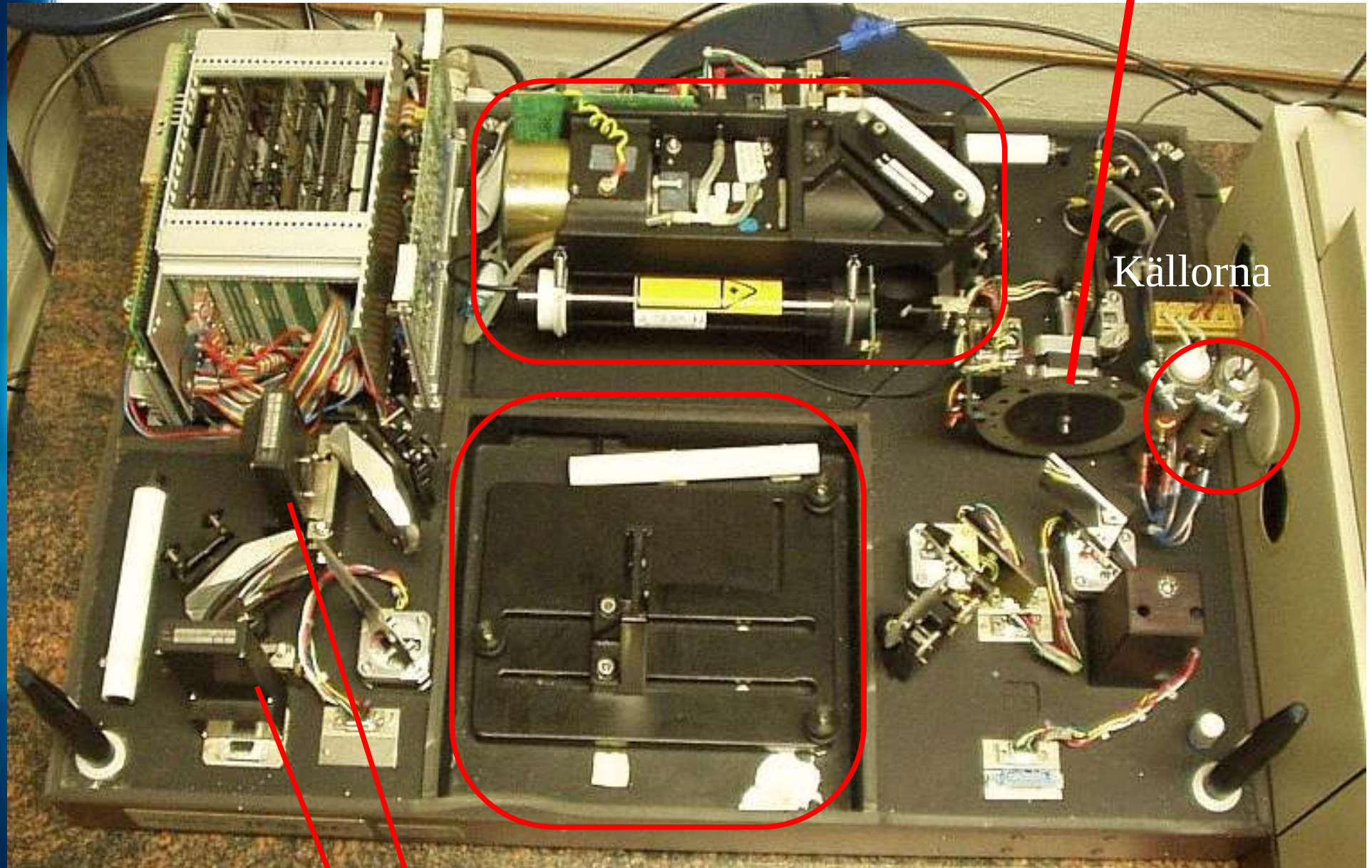
Interferometer

Jaquinotöppning

Källorna

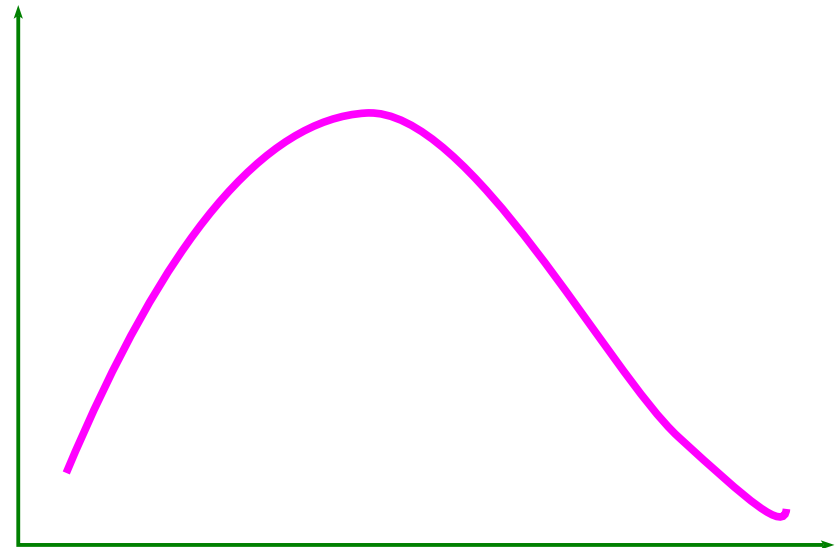
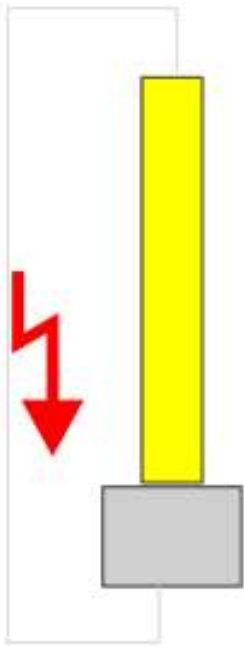
Detektorerna

Provkammare



Ljuskällan

Nernstkälla

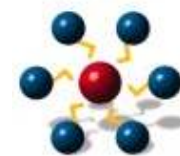
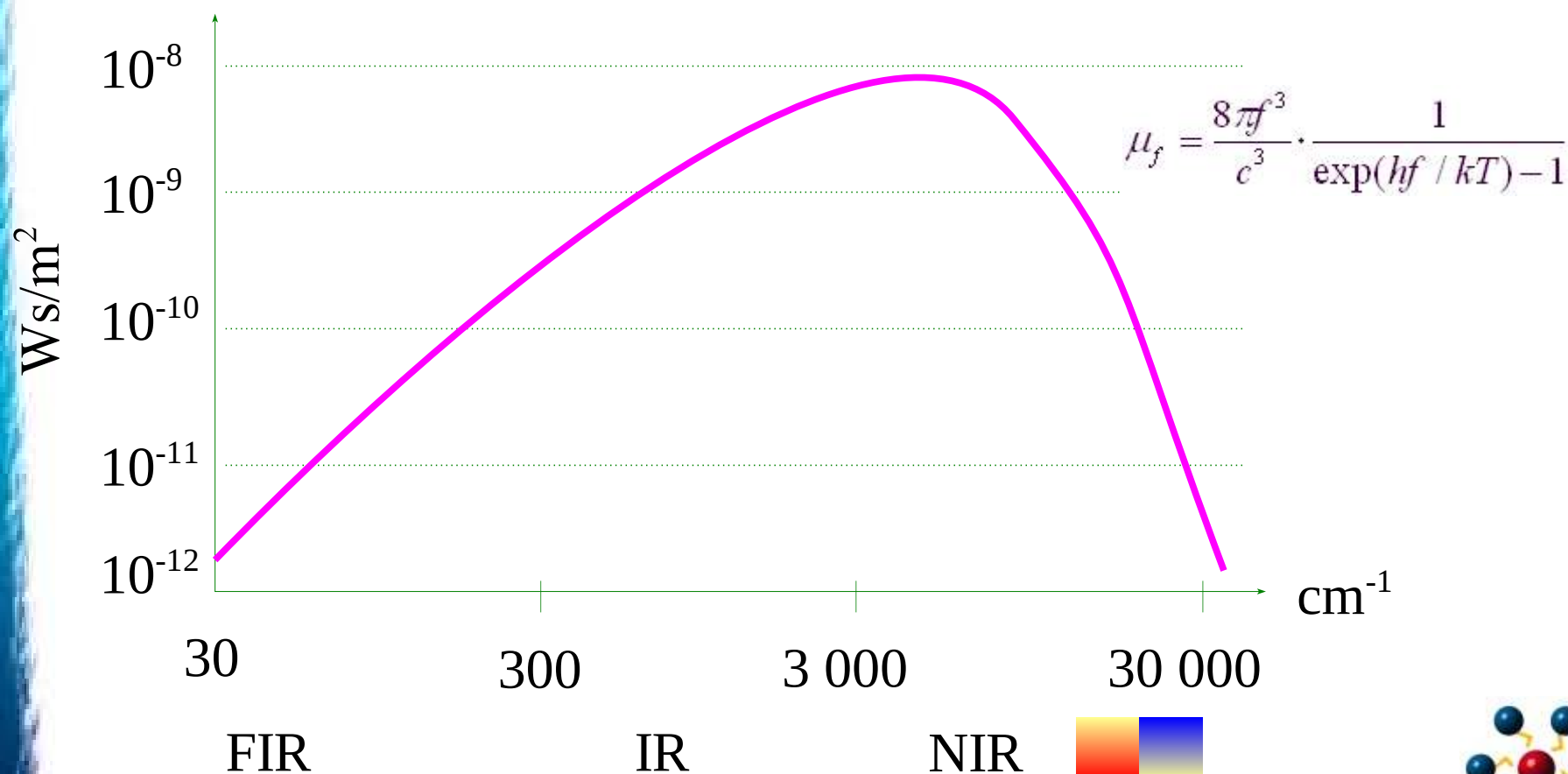


Sällsynta jordarters oxider.
Ca. 1500 °C



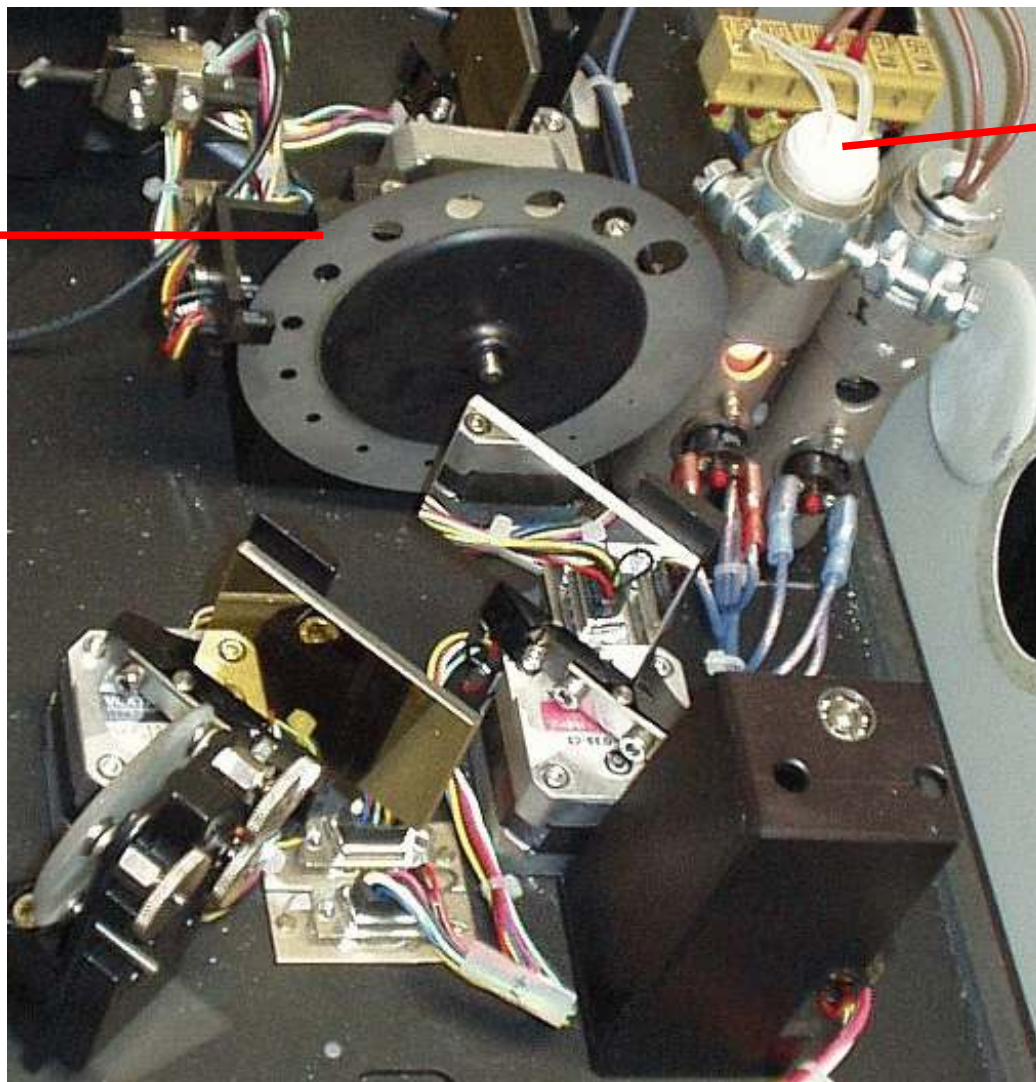
Källans ljusintensitet

Glödande svarta kroppens strålning

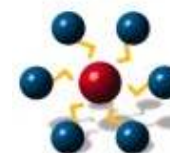


Källan

Jaquinots
öppning

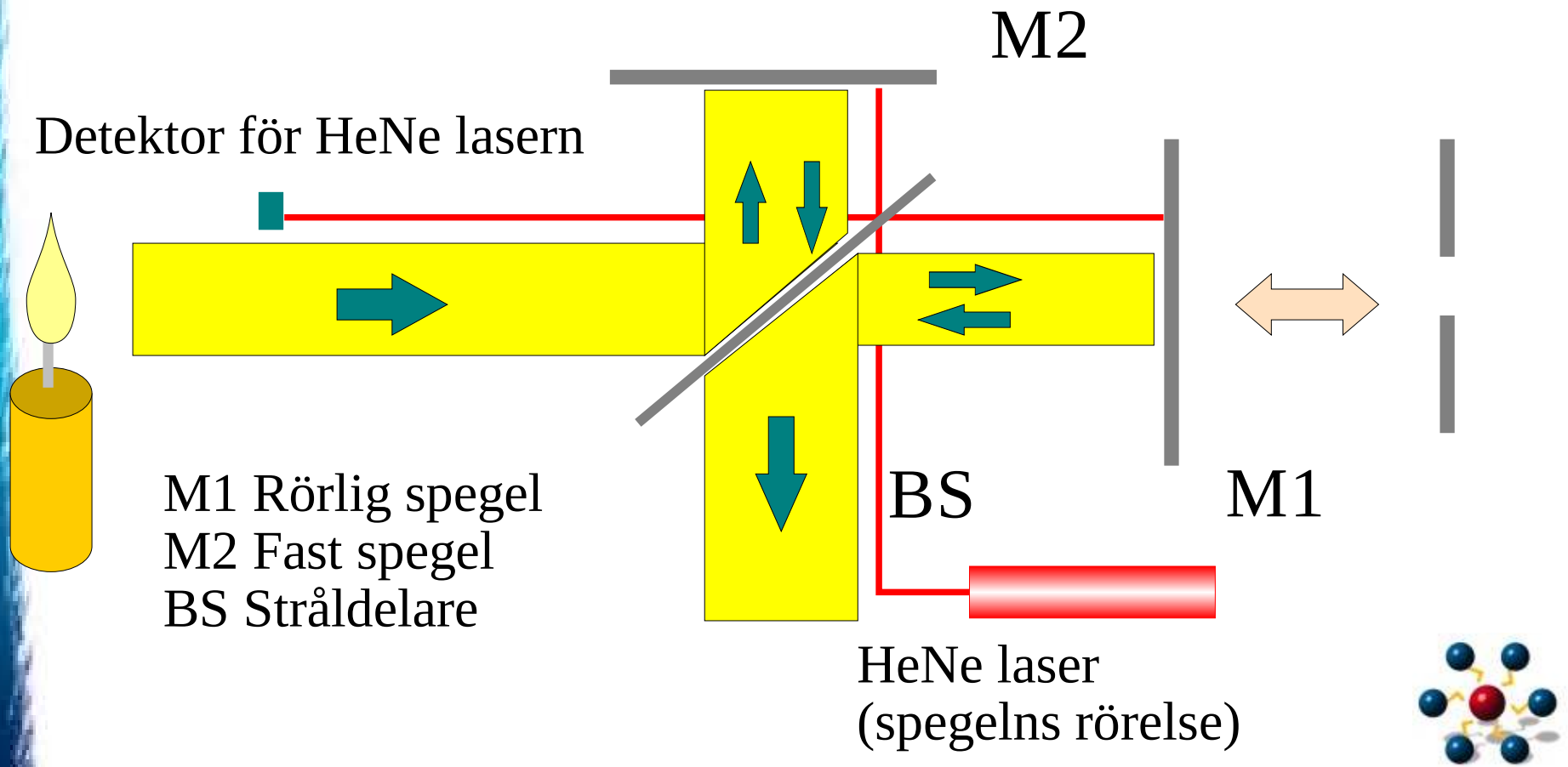


Nernstkälla



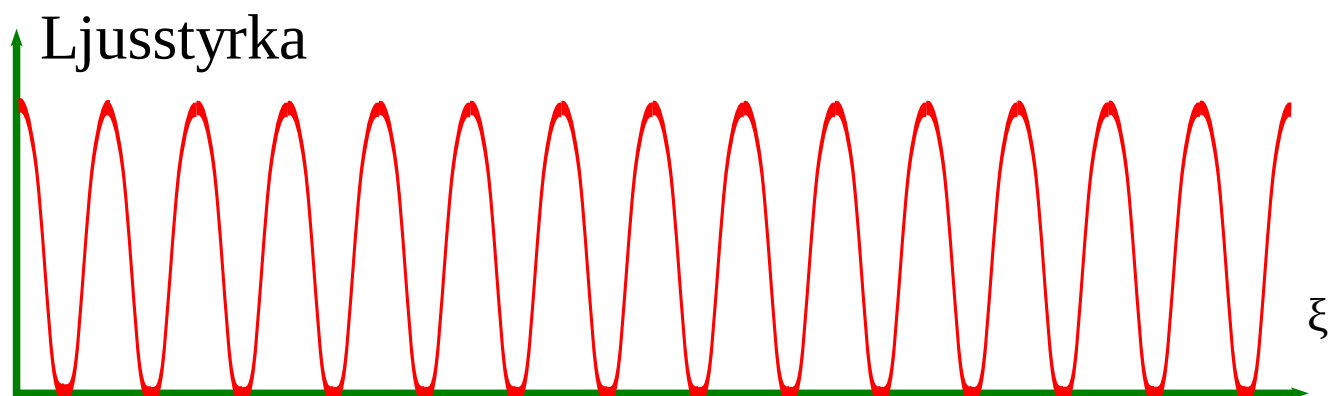
Michelsons interferometer

Schematiskt

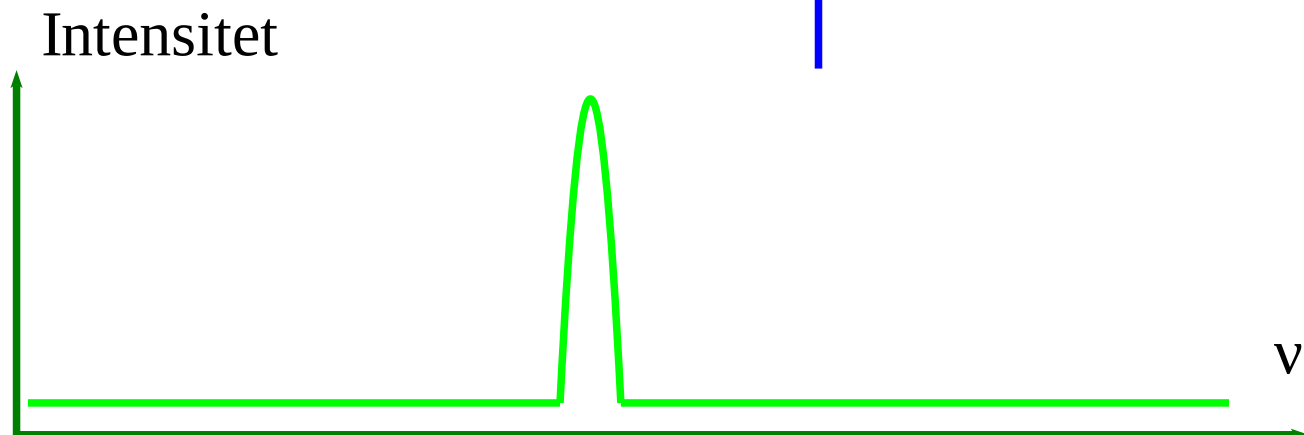


Fouriertransformation

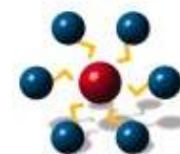
Interferogram



Spektrum



FT



Fouriertransformation

$$S(\nu) = \int_0^{\infty} P(\xi) \cos(2\pi\nu\xi) d\xi$$

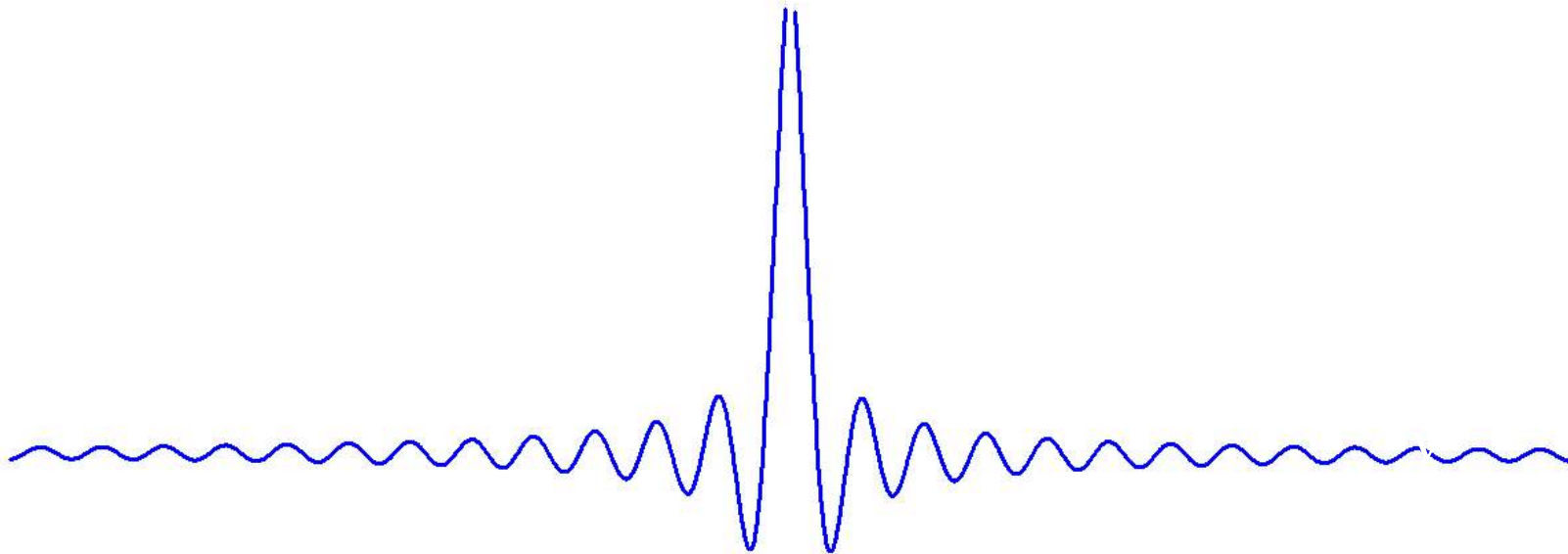
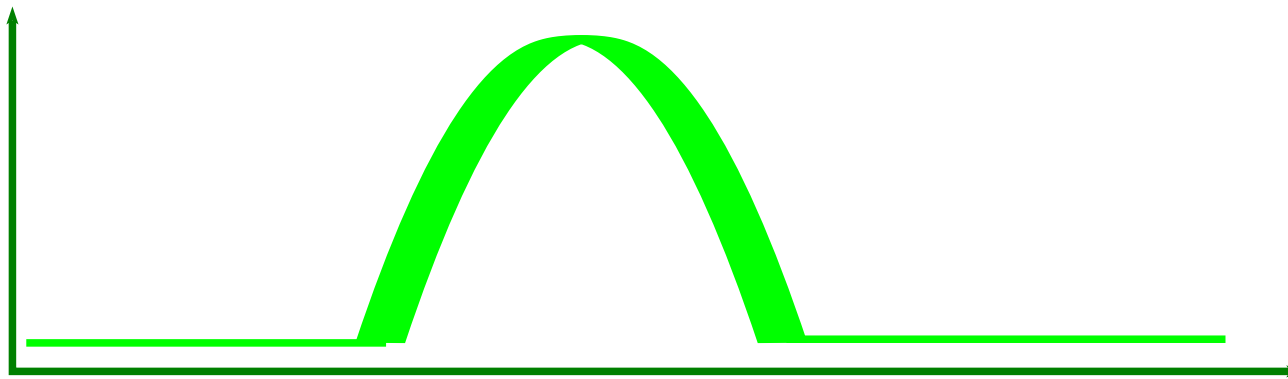
IR-
spektrum

Interferogram



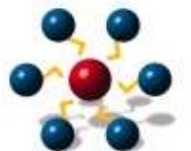
Fouriertransformation

Brett spektrum
Intensitet



Fouriertransformation

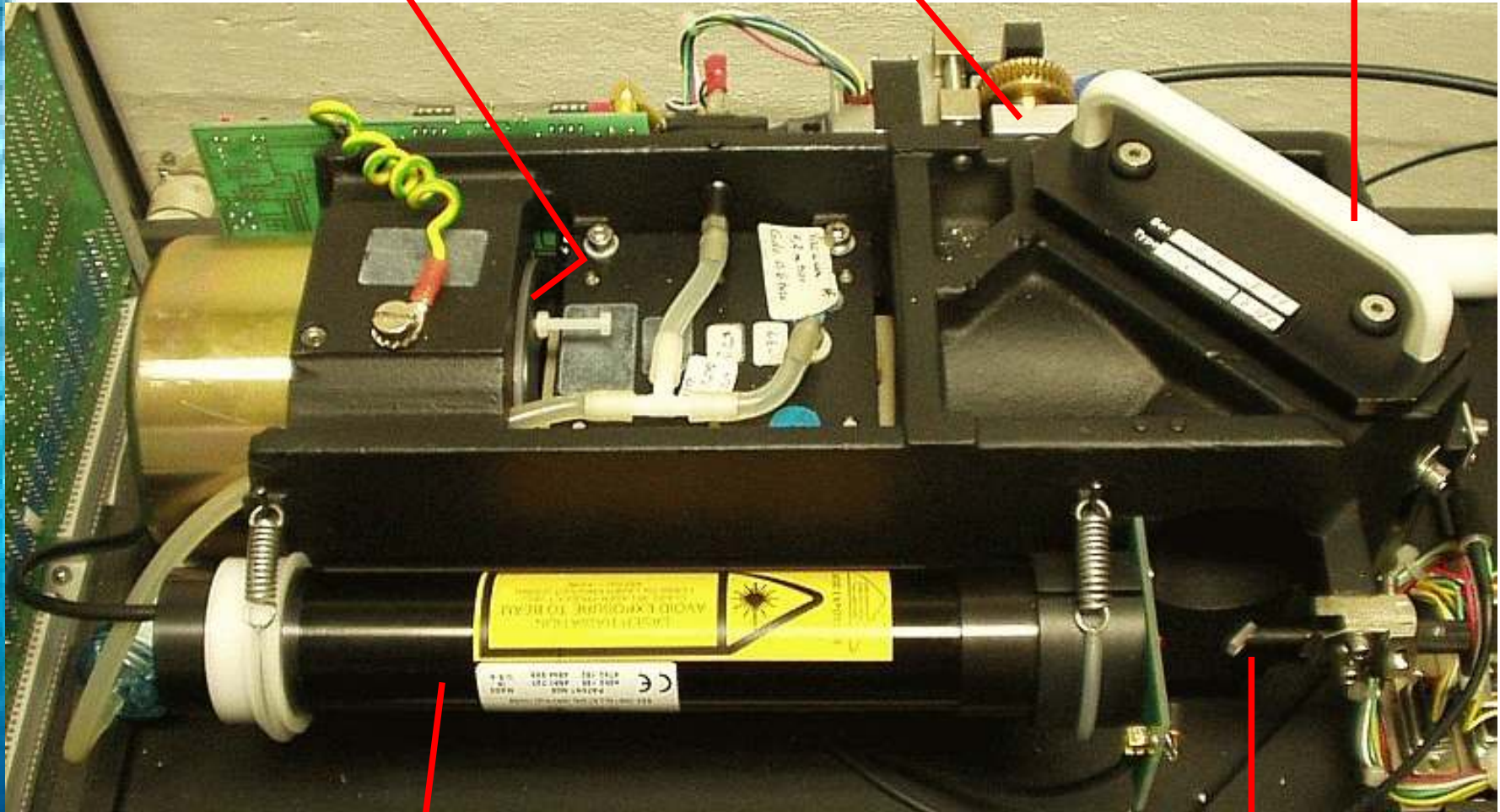
- ⇒ Längre spegelrörelse $\Rightarrow$ högre upplösning
 - ⇒ Tätare provtagning $\Rightarrow$ högre frekvens
-



Rörlig spegel

Fast spegel

Stråldelare



HeNe laser

Spegel för laserstrålen



Interferometer

Stråldelare



Interferometer

Stråldelare

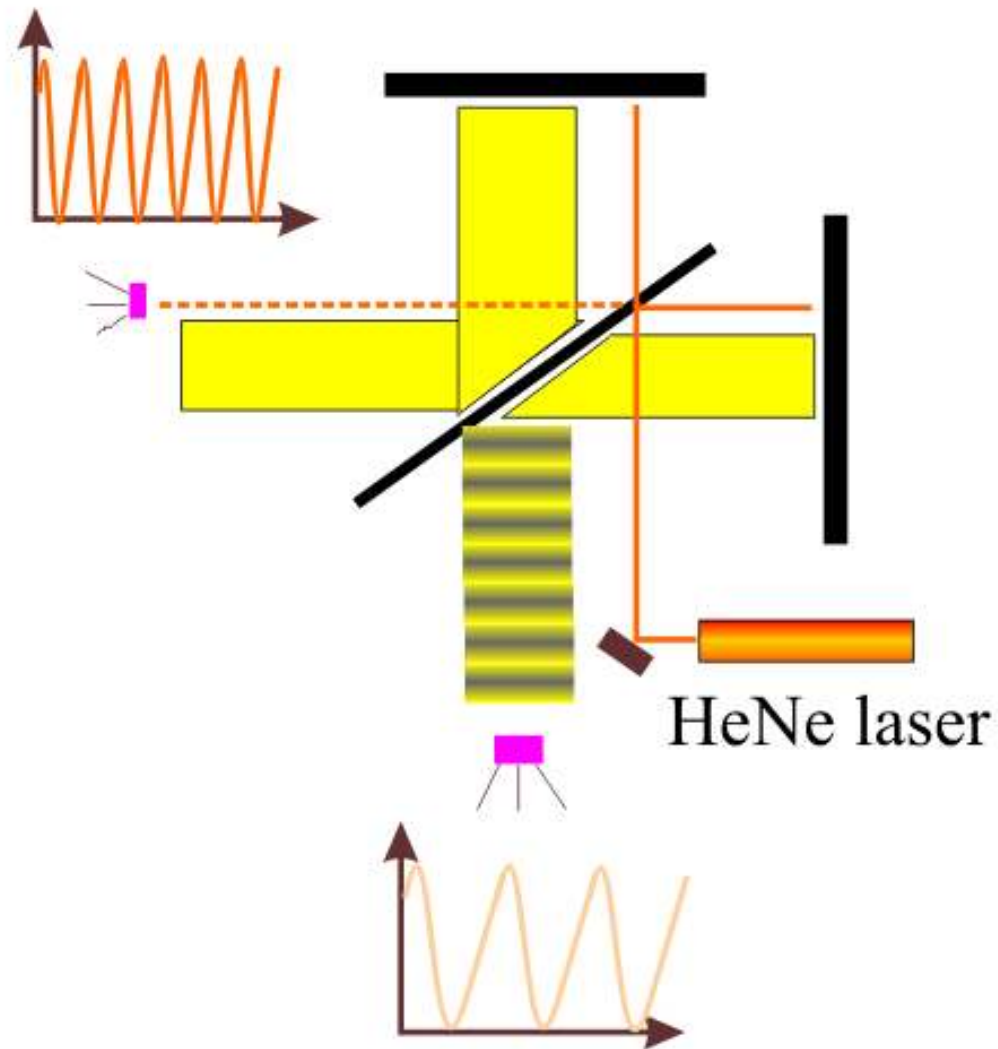


Interferometer

Stråldelare



Interferometers funktion



Interferogram

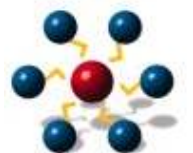
Central burst

I

ξ = spegelns rörelse

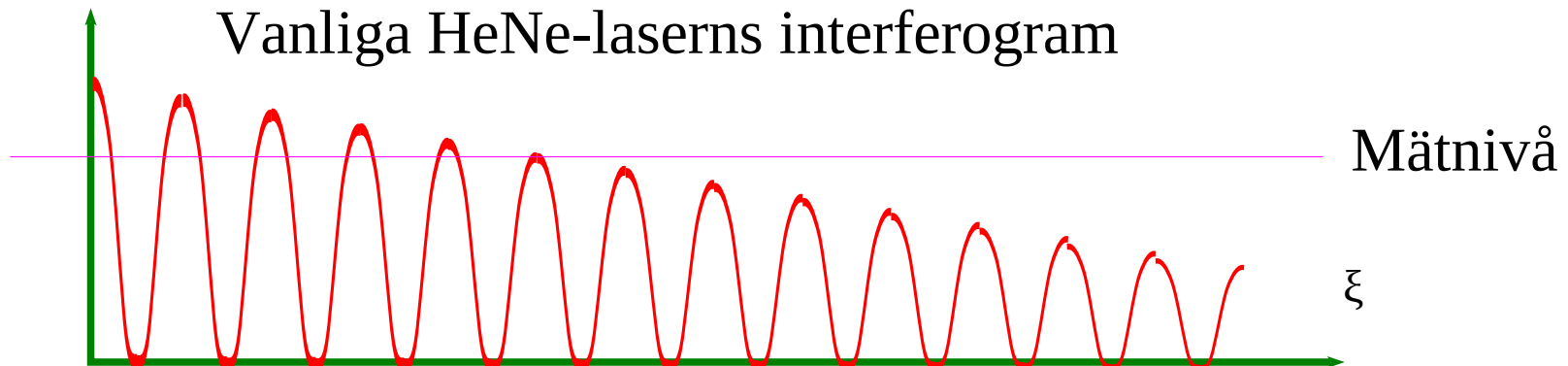
Mätområde

ξ

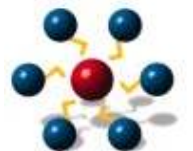
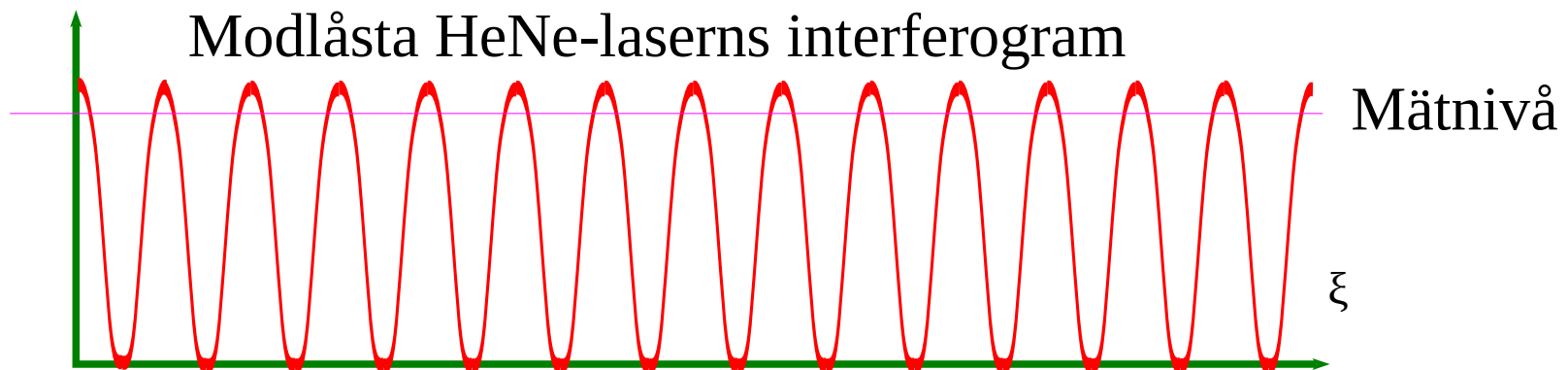


Laser för avståndsmätning

Vanliga HeNe-lasers interferogram



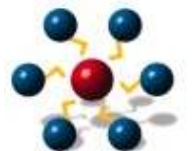
Modlåsta HeNe-lasers interferogram



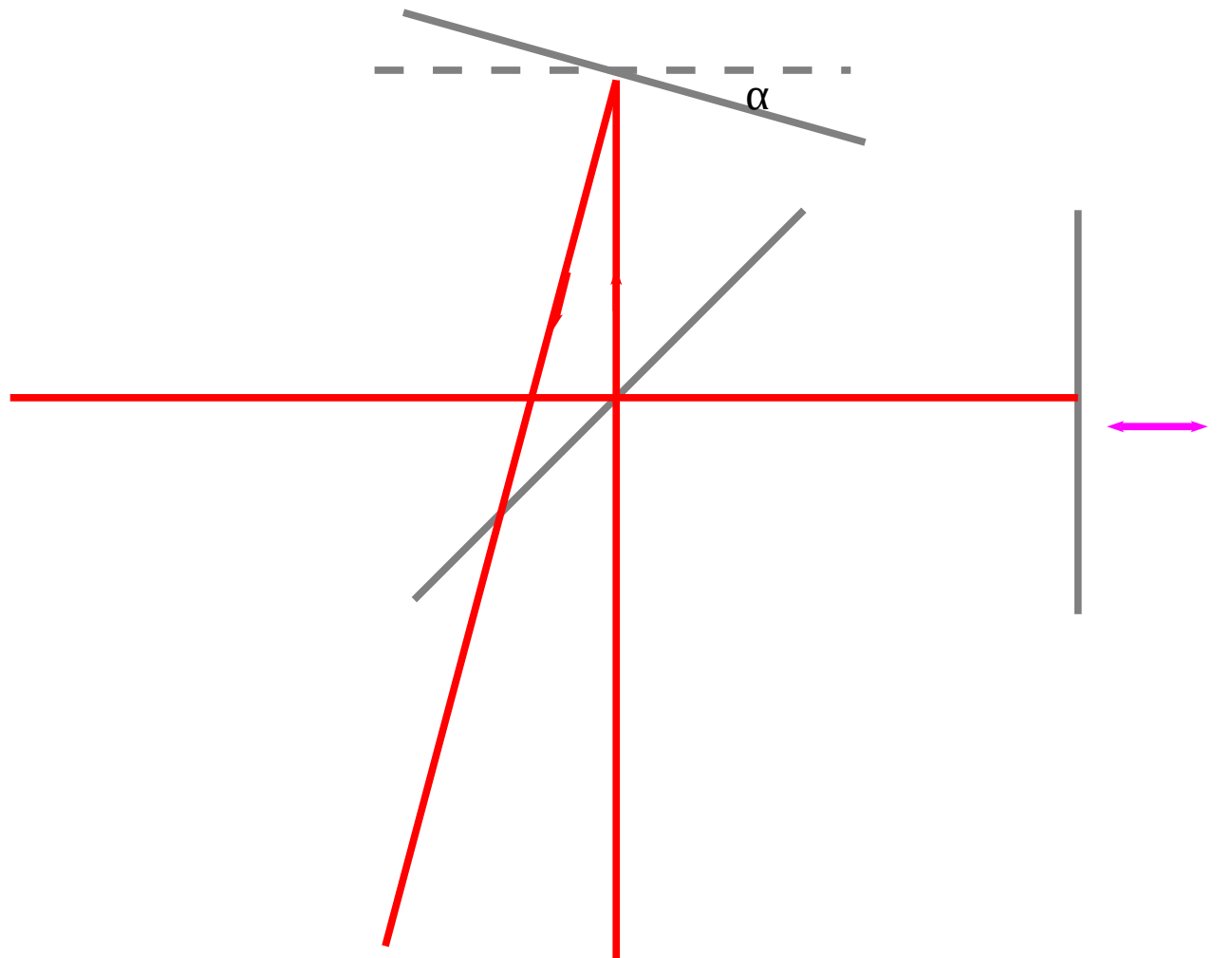
Spektrometers linjering

Linjering betyder, att interferometers speglar, speciellt den fasta spegeln, vrids så, att det centrala bandet (central burst) har den största möjliga ljusstyrkan.

Då interferometern inte är rätt justerad, lider spektret av dålig känslighet och många andra brister.



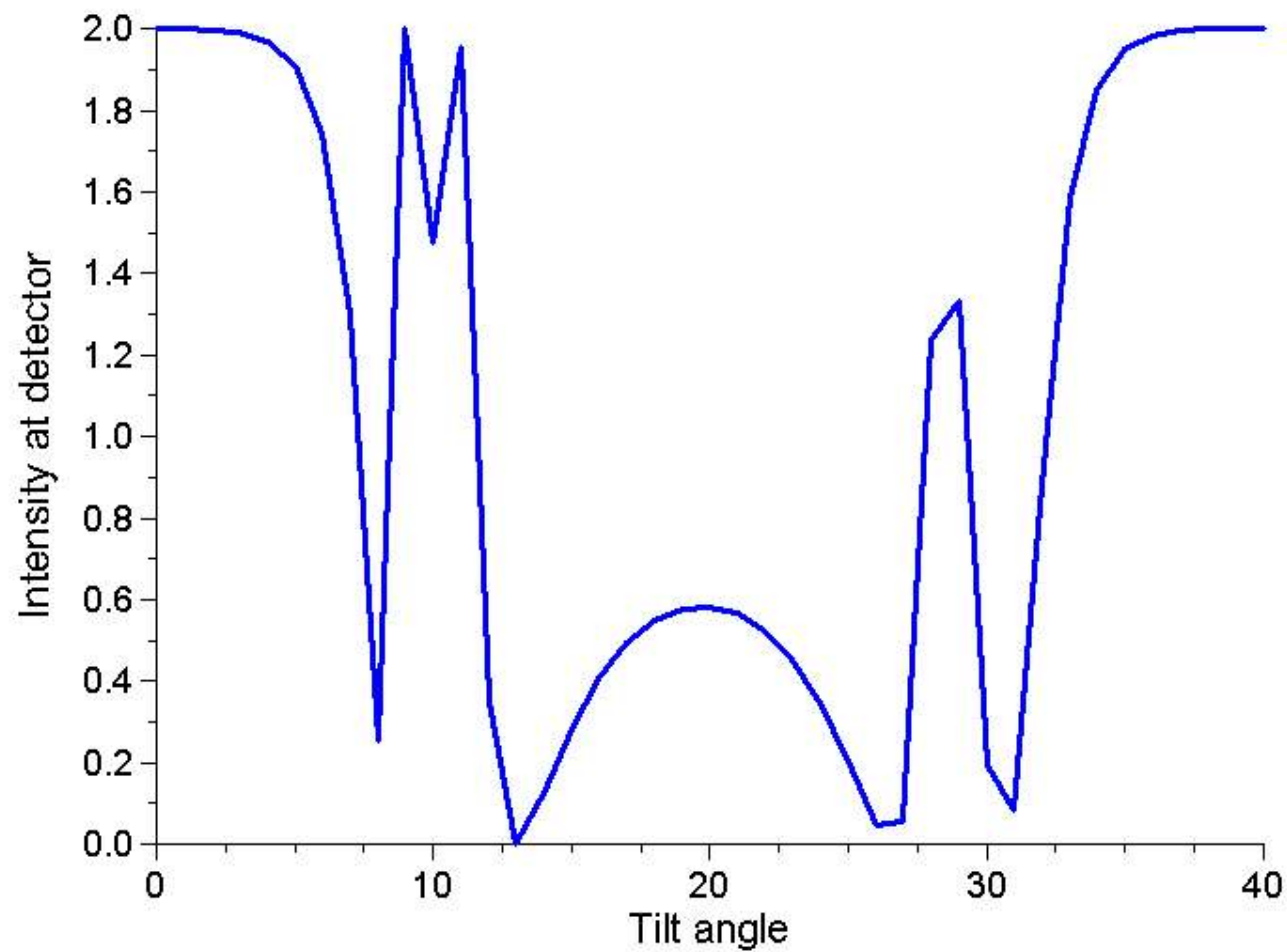
Fel linjering



Fel linjering minskar upplösningen och förvränger vågtalsaxeln.



Fel linjering



Detektor

Halvledardetektortyperna

⇒ DTGS

- ▶ Deuterium TriGlycine Sulphate
- ▶ Fungerar vid rumstemperatur
- ▶ Brett vågtalsområde
- ▶ Inte speciellt känslig

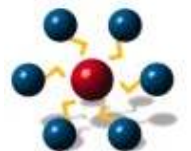
⇒ MCT

- ▶ Mercury Cadmium Telluride
- ▶ Känslig
- ▶ Måste kylas
- ▶ Snävt vågtalsområde



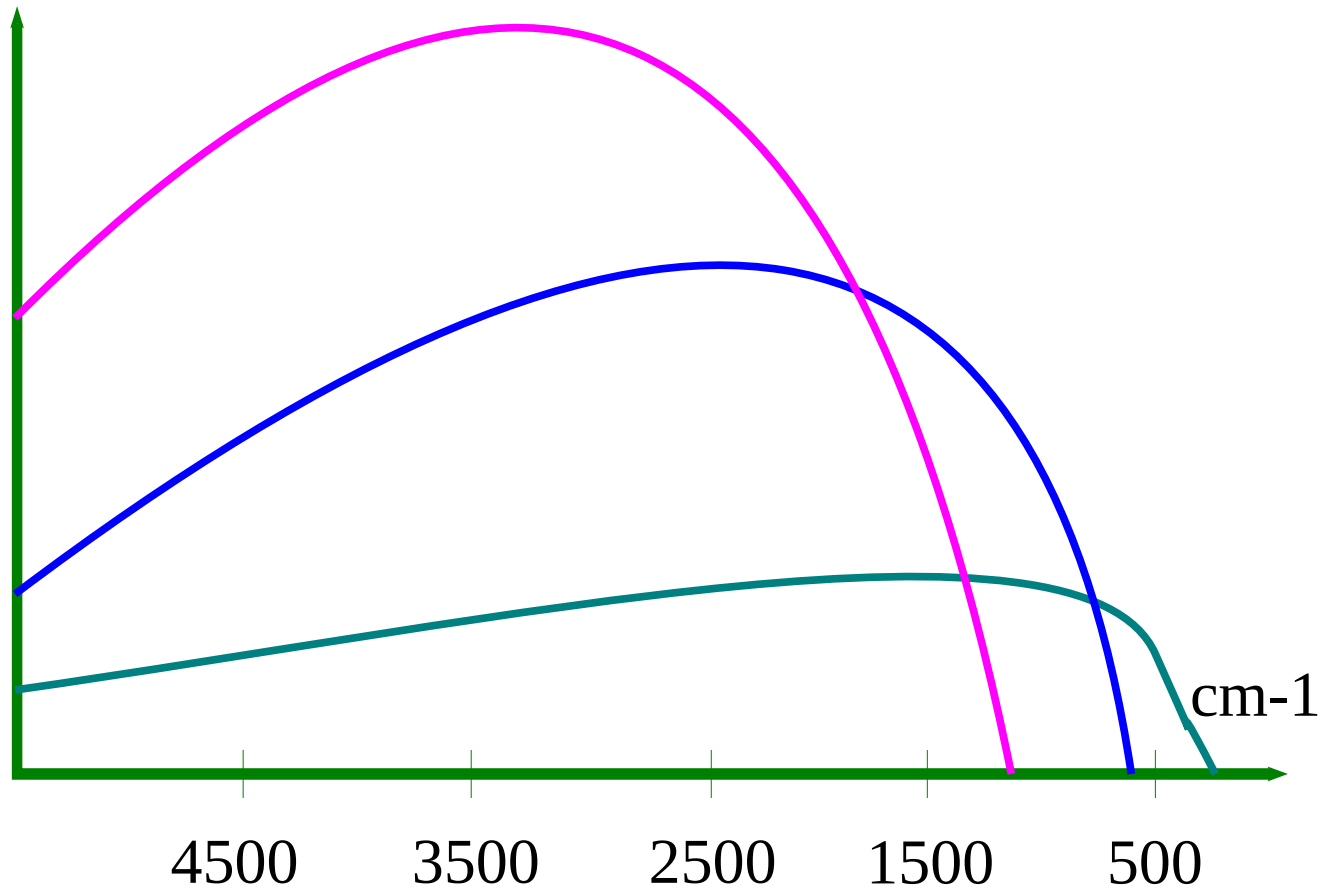
IFS66

Mid-IR detektor



MCT-detektorerna

Känslighetskurvor



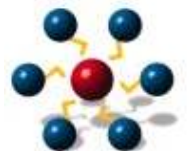
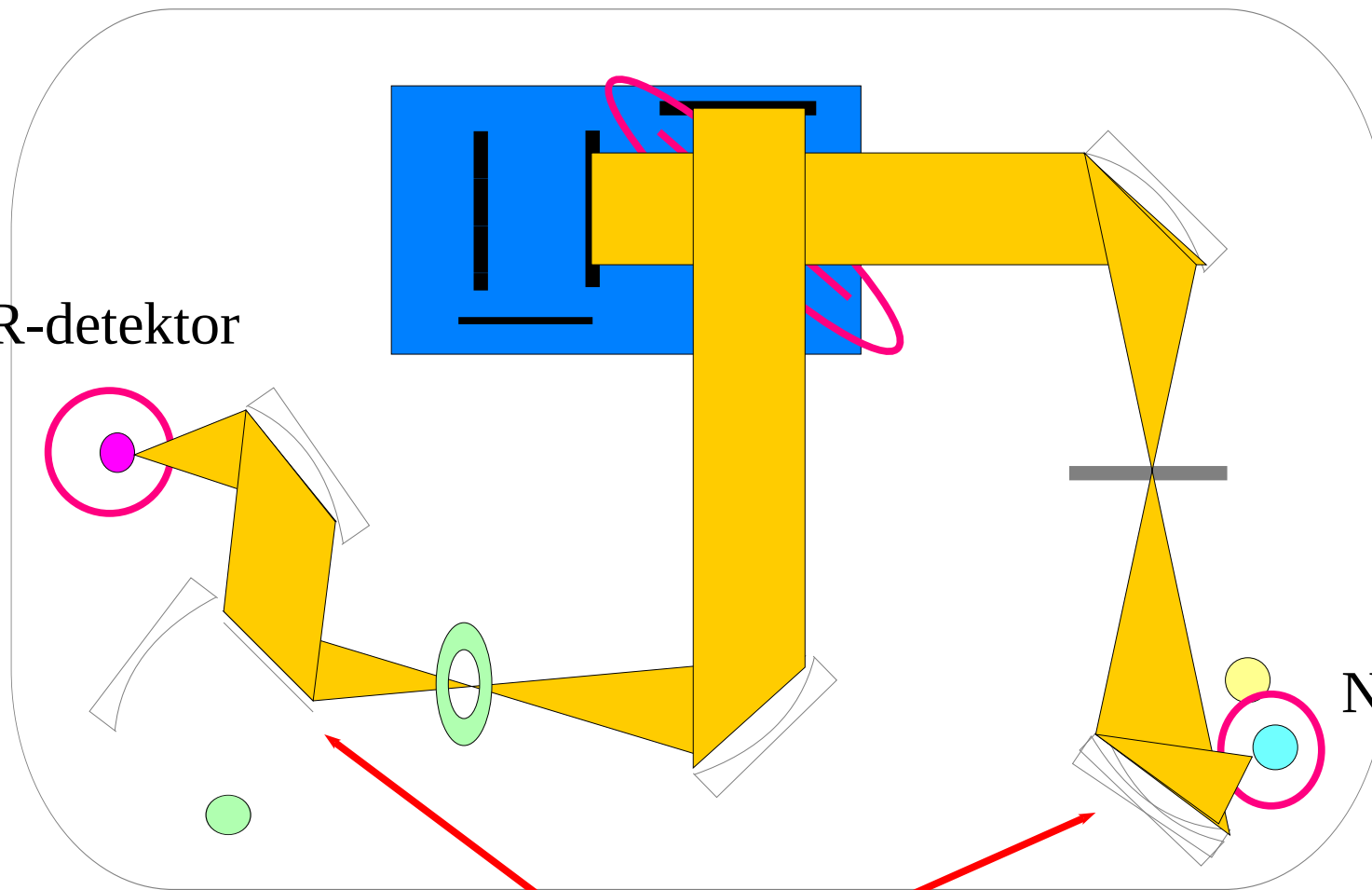
NIR-option

NIR-stråldelare

NIR-detektor

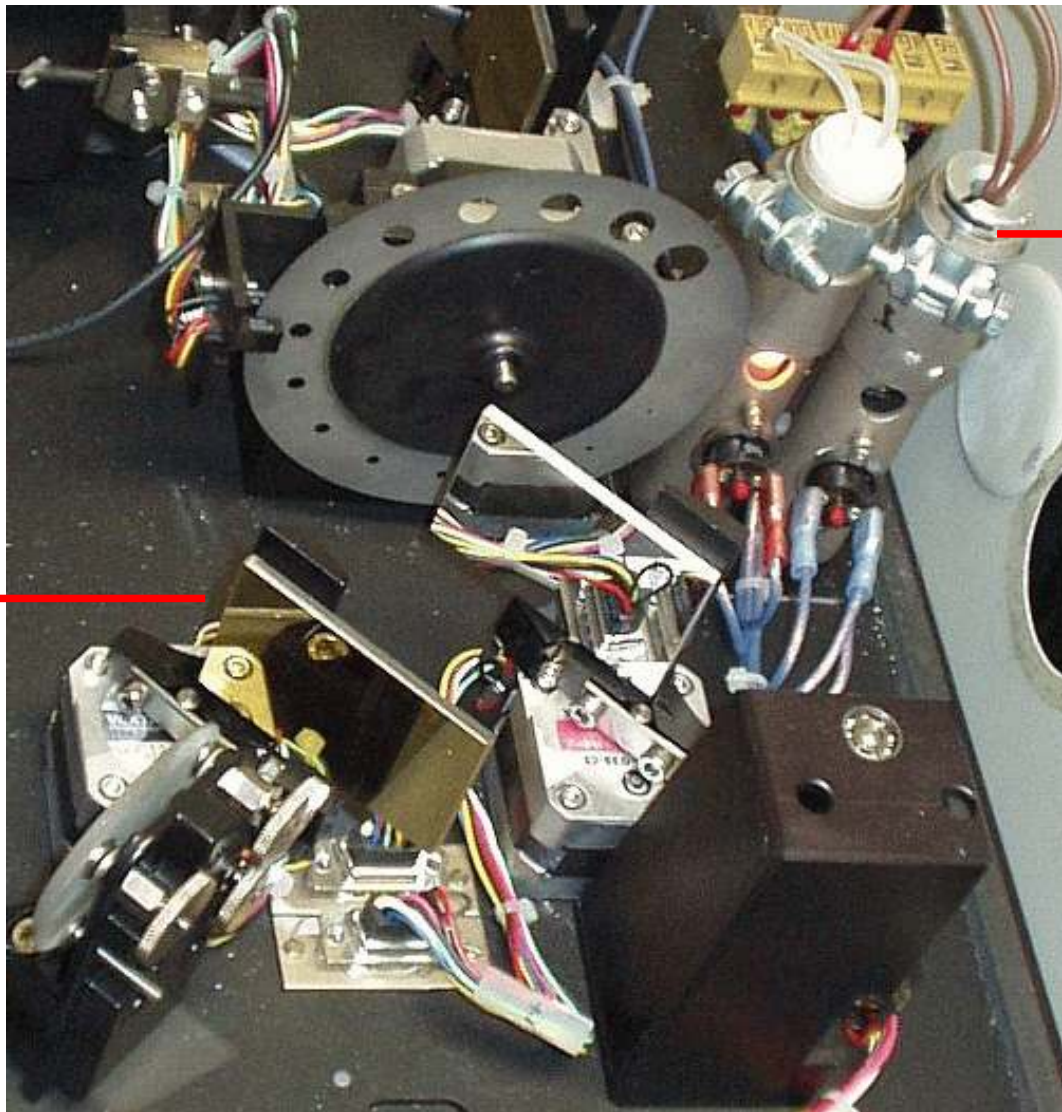
NIR-källa

Vridbara speglar



IFS66

Strålens
styrning



NIR-källa



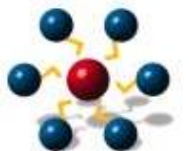
NIR-källa

⇒ Wolframglödlampa



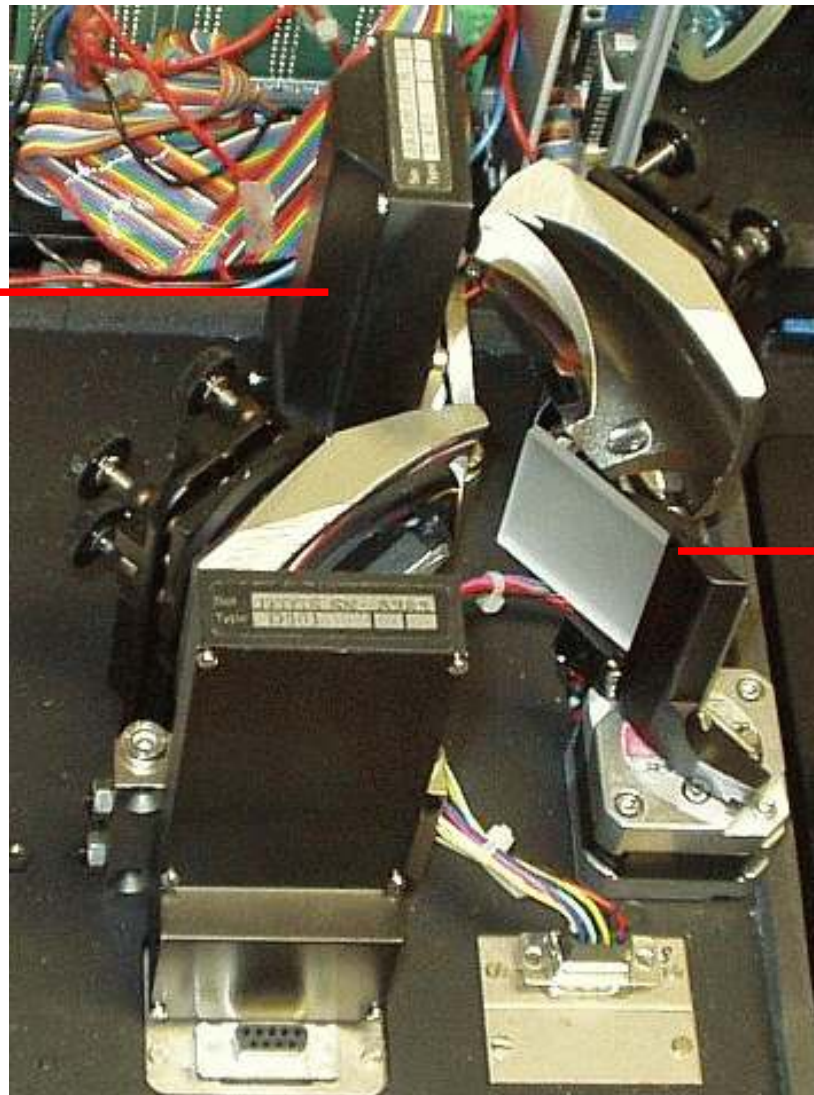
NIR-stråldelare

⇒ Oftast inbyggd i apparaten (i en revolver)



IFS66

NIR-detektor

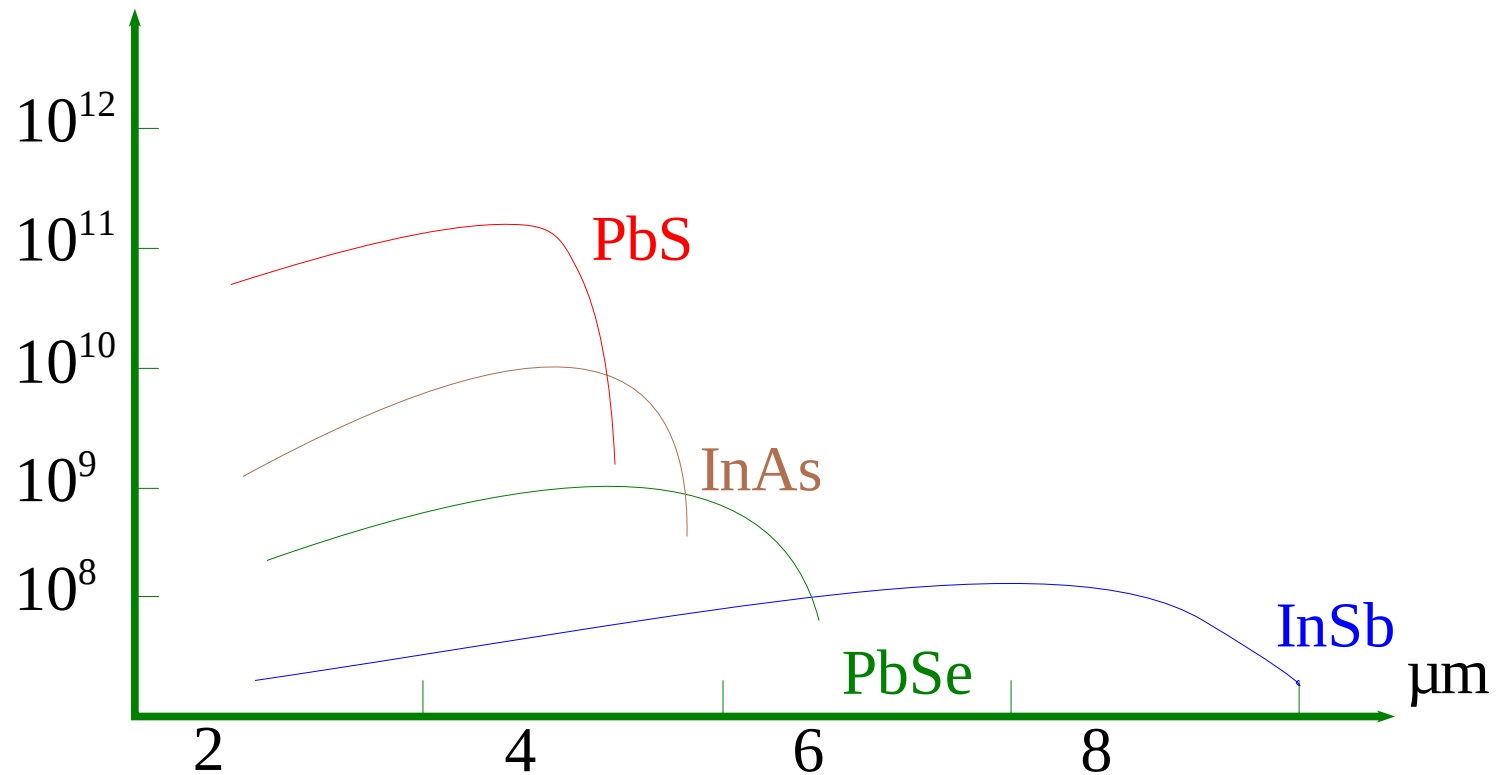


Strålens styrning

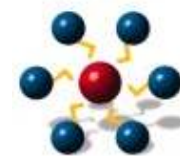
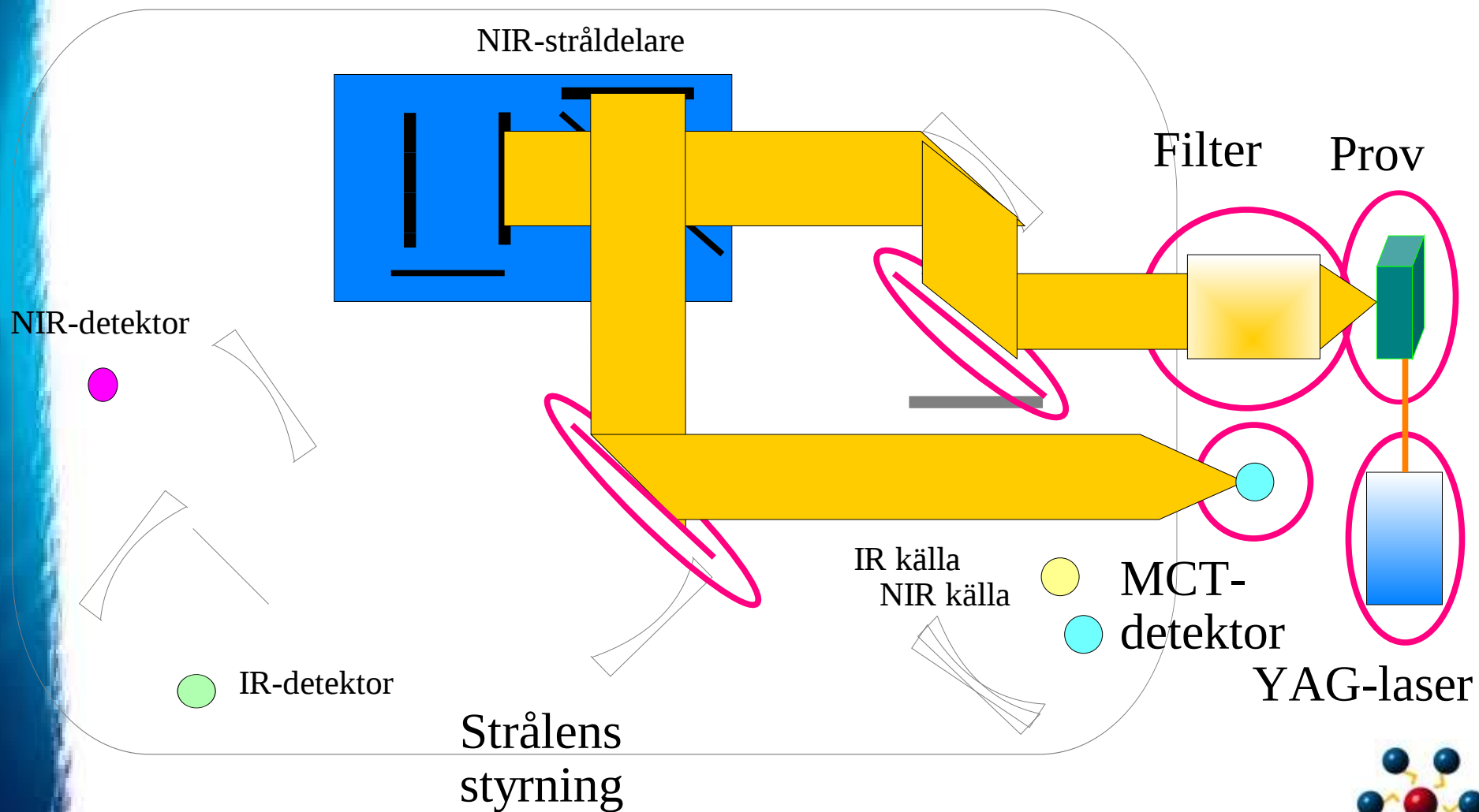


NIR-detektor

⇒ Diioddetektorer

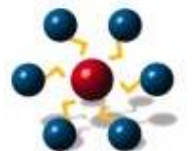


Ramanoption



Väckning

- ⇒ Normalt en laser
 - ▶ YAG-laser: NIR-ljus, $\lambda = 1 \mu\text{m}$
 - ▶ Ibland laser som emitterar synligt ljus, t.ex. HeNe eller Ar
 - ▶ Hög ljusstyrka => stark Raman-signal
 - ▶ Lätt att fokusera i provet



Filter

